

Strenghts	Weakness
Top rank scientific level	Words like "paramount", "epochal", "unique", "compelling", "historical", "enormous", "extraordinary", "invaluable". "dramatically" and so on are so obvious in this context to become upsetting
This is a well-written, timely, proposal of paramount importance. The science goals are extremely appealing and can have a profound impact on a wide range of modern astrophysics topic. Effort should be made to prioritize this kind of observations, multi-messenger astrophysics just started and the scientific return will be enormous. The team is large and strong, and has a consolidated experience in working with REM and previous results are impressive and have been published on one of most important scientific journal worldwide. The proposal make a very good use of the REM capabilities, the observing strategy is well presented and thought and the likelihood of achieving the science goal looks consequently high.	None
The proposal aims at helping in the search of optical-nIR events associated to GWs in relatively bright host galaxies. This activity will be carried out by the GRAWITA team that demonstrated its crucial and leading role in the discovery and characterization of the EM counterpart of GW170817. The project will uncover the O3 LVC run that will start in early 2019. Definitively to be persued.	None
Proposal focusing on one of the most interesting and hot topics in astrophysics. Based on lessons learnt from the successful case of GW170814.	None
Follow-up of gravitational wave sources is a promising field, with a huge discovery potential. Multi-color observations are a critical first step in the identification of the true GW counterpart. IR coverage is particularly important to characterize the possible kilonova emission from the merger ejecta. This is a strong team with access to a wide array of other facilities to leverage the REM results.	The proposal does not discuss how the galaxy strategy will be implemented, and how candidates will be vetted (e.g. human inspection, automatic pipeline, or both). Additional details on the rapid data reduction and classification would have made the proposal stronger.

The proposal is well motivated and focuses on a hot topic in a rapidly changing field.

The authors do not comment on the scientific (or strategic) outcome of previous observing time, except of course the observations of the EM counterpart of GW170817 during AO35.

The need of observing with the two instruments to fully characterize the SED, although in some way obvious, is not discussed at length (e.g., the authors do not discuss what could be done with a lower number of bands).
